

Degeneracy vs. Energy Level Scaling for Hydrogen

Sabrina Gonzalez Pasterski

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I explore the relationship between degeneracy and energy for the Hydrogen atom.

The energy levels of the Hydrogen atom from quantum mechanics are given by:

$$E_n = -\frac{E_0}{n^2} \quad E_0 = \frac{me^4}{2\hbar^2}, \quad n \in \mathbb{Z}^+ \quad (1)$$

in c.g.s. units, where no perturbations are included. The degeneracy of each level is n^2 . If it were possible to ignore electron-electron interactions and fill the energy levels following the Pauli Exclusion Principle, a total of $2n^2$ electrons could occupy each E_n , corresponding to one spin-up and one spin-down electron in each of the n^2 states. To make the atom neutral when it has a fixed number of electrons, the charge of the nucleus could be increased, scaling E_0 . In this manner, a completely filled shell would contribute $-2E_0$ to the total energy of the system, regardless of the n that the shell corresponds to.

This is consistent with the Hellman-Feynman theorem:

$$\frac{\partial E_n}{\partial e} = \left\langle \psi_n \left| \frac{\partial H}{\partial e} \right| \psi_n \right\rangle \rightarrow \left\langle \frac{1}{r} \right\rangle = \frac{1}{a_0 n^2} \quad (2)$$

where $a_0 = \frac{\hbar^2}{m_e e^2}$ is the Bohr radius. Higher energy shells of Hydrogen are further away from the proton. Increasing the amount of orbiting electron charge by n^2 would cancel this effect and give the same potential energy for each shell.

The total energy of the system is linear in the filling fraction ν , representing the number of filled energy levels. Each electron within a given shell contributes equally, while a totally filled shell contributes $-2E_0$. Figure 1 shows how the total energy decreases with ν while the density of allowed values of ν increases as $2n^2$ for shells that can hold more electrons. In the ground state, at $T = 0$, the lowest energy levels are filled.

The symmetry of a constant $D_n E_n$, for degeneracy D_n , also effects the relative probability for finding a single electron in an n -shell versus m -shell state. No electron-electron interactions are being ignored, although only relative probabilities are studied since the sum over all levels diverges. If the probability distribution for the state of a single particle can be treated with Maxwell-Boltzmann statistics, then:

$$\frac{\mathbb{P}(n)}{\mathbb{P}(1)} = n^2 e^{\frac{E_0}{k_B T} \left(\frac{1}{n^2} - 1 \right)}. \quad (3)$$

In the high temperature limit, $\mathbb{P}(n) \propto n^2$, so that the average contribution to the energy of the system: $\mathbb{P}(n)E_n$ is independent of n .

Equation 3 also shows that $\mathbb{P}(n)$ has a minimum at $n = \sqrt{\frac{E_0}{k_B T}}$, corresponding to $k_B T = \frac{E_0}{n^2} = |E_n|$. When the characteristic thermal energy $k_B T$ is closest to E_n , the n^{th} energy level is least likely to be populated. This relates to stimulated absorption since the amount of thermal energy is just enough to ionize the n^{th} energy level. For instance, a gas of hydrogen atoms (as opposed to H_2) at the temperature of the sun's photosphere would have a minimum relative population for the $n = 5$ energy level according to this Boltzmann distribution. Higher temperatures preferentially depopulate shells with smaller n .

If instead, we had a constant $D_n E_n$ with $D_n \propto n^j$ rather than n^2 , the least populated state would still correspond to $k_B T = |E_n|$. If the power of the degeneracy is different, such that $D_n \propto n^f$, $E_n \propto n^{-j}$, the minimum population occurs at $k_B T = \frac{j}{f} |E_n|$. A similar extremum occurs for energy levels that are positive and increase with n , reminiscent of Rotational Raman Scattering.

Since degeneracies broken by perturbations usually only cause small splittings in the energy levels of a system, one could imagine that there would be cases where the $\frac{j}{f}$ proportionality factor could be used to define an effective order for the degeneracy of energy levels.

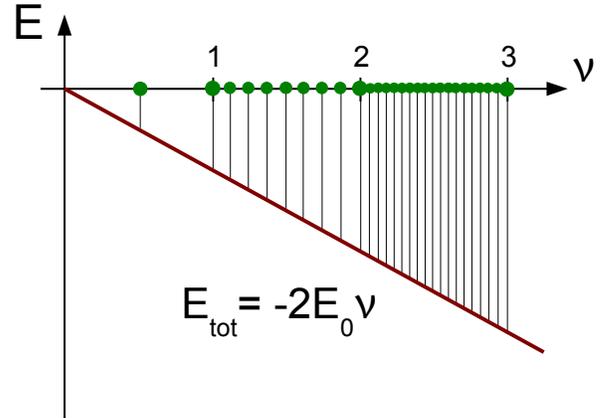


FIG. 1. Total energy as a function of the filling fraction for the ground state of an idealized multi-electron atom.